#### ISABEL INC Model for High-Energy Hadron-Nucleus Reactions

SOREQ

#### Y.Yariv

#### SOREQ NRC, Yavne 81800, Israel

International Topical Meeting on Nuclear Research, Application and Utilization of Accelerators **Satellite Meeting on Spallation Reactions** IAEA, Vienna, Austria 4-8 May, 2009







#### איזבל ISABEL

אשד תוך גרעיני Eshed Toch Gar'ini → ETGAR אתגר Etgar = Challenge

SOREQ
$\sim$

#### History

- M.L.Goldberger, Phys. Rev. 74, 1269 (1948)
- N.Metropolis et al., Phys. Rev. 110, 185 (1958); Phys. Rev. 110, 204 (1958)
- VEGAS: K.Chen et al., Phys. Rev. 166, 949 (1968)
- ISOBAR: G.D.Harp et al., Phys. Rev. C8, 581 (1973); C10 2387 (1974)
- ISABEL: Y.Yariv and Z.Fraenkel, Phys. Rev. C20, 2227 (1979); Phys. Rev. C24, 488 (1981)
- ♦ ETGAR...



#### **Nuclear Model**

- Continuous charge distribution Folded Yukawa.
   Nucleus divided into several regions of constant density.
   Ratio of proton to neutron density Z/(A-Z)
- Momentum distribution degenerate Fermi Gas  $E_{F_i} = (\hbar^2 / 2m_i)(3\pi^2 \rho_i)^{2/3}$

 $i = proton, neutron; m_i = nucleon \_mass; \rho_i = density$ 

Potential depth (J.N. Ginocchio, Phys. Rev. C17, 195 (1978))

$$V_{i} = E_{F_{i}} + (Separation \_ Energy)_{i}$$
$$V_{\Delta^{++}} = V_{p}; V_{\Delta^{+}} = V_{p} + \frac{(V_{p} + V_{n})}{3}; V_{\Delta^{0}} + \frac{(V_{p} + V_{n})}{3} = V_{n}; V_{\Delta^{-}} = V_{n}$$



#### Hadron-Hadron Cross Sections (1)

#### **⊗ N+N**

- σ<sub>tot</sub>, σ<sub>inel</sub>, σ<sub>el</sub>
   G.D.Harp, Phys. Rev. C10, 2387 (1974)
   Arndt phase shift analysis
- dσ<sub>el</sub>/dω
   P.C.Clements, L.Winsberg, UCRL 9043 (1960), unpublished



#### Hadron-Hadron Cross Sections (2)

#### $\circledast N{+}N \to N{+}\Delta$

 Type of outgoing N, Δ determined by Isotopic Spin consideration
 Z.Fraenkel, Phys. Rev. 130, 2407 (1963)

• Mass of  $\Delta$  is chosen from distribution:

$$P(m_{\Delta}, E_{cm}^{N+N}) = const. * \sigma_{tot}^{\pi^{+}+p}(E_{cm}^{N+N}) * F(m_{\Delta}, E_{cm}^{N+N})$$

 $m_{\pi} + m_N < m_{\Delta} < m_{\pi} + m_N + 500 MeV$ 

F = two body phase factor for the produced N+ $\Delta$ S.Lindenbaum and R. Sternheimer, Phys. Rev. **105**, 1874 (1957); **109**, 1723 (1958); **123**, 333 (1961)



#### Hadron-Hadron Cross Sections (3)

#### $\Delta + N \rightarrow N + N (\pi capture)$

- Type of outgoing N, Δ determined by Isotopic Spin consideration
- σ, dσ/dω calculated from inverse process (Δ production) using the principle of "detailed balance"

7

•  $\Delta$  production calculated using theoretical model (OPE)



#### Hadron-Hadron Cross Sections (4)

#### $\otimes \Delta + N \rightarrow \Delta' + N'$ ("exchange")

Naively two step process:

Decay of initial Isobar,  $\Delta \rightarrow \pi + N'$ 

Interaction of decay  $\pi$  with another Nucleon,  $\pi + N \rightarrow \Delta'$ 

G.D.Harp et al., Phys. Rev. C6, 581 (1973),
Z.Fraenkel, Nuovo Cimento 30, 512 (1963)
Z.Fraenkel, Phys.Rev. 130, 2407 (1963)





#### Hadron-Hadron Cross Sections (5)

# $\label{eq:phi} \begin{array}{l} & & \pi + N \rightarrow \Delta \rightarrow \pi' + N' \\ & \mbox{(elastic \& charge exchange)} \end{array}$

- Experimental dσ/dω + isospin considerations
   G.Giacomelli et al., CERN/HERA 69-1 (1969)
- For Δ decaying without interaction proper π+N differential cross section
- Solution Isotropic  $\Delta$  decay after scattering or exchange



#### Hadron-Hadron Cross Sections (6)

#### ${\circledast} \Delta \to \pi{\textbf{+}} N$

Energy dependant Δ width
 J.N. Ginocchio, Phys. Rev. C17, 195 (1978)



#### 

- Fast rearrangement: ρ<sub>i</sub> of the "partner type" Fermi sea is uniformly reduced for the whole nucleus
- Slow rearangement: "partner type" hole of radius r is punched in the position of the interaction. No interactions are allowed in the hole with particles of "partner type".



## **Pauli Blocking**

#### Options:

- Full Pauli Blocking: Interaction resulting in nucleon falling below Fermi sea is forbidden
- "Depleted" Pauli Blocking: Reaction resulting in nucleon falling below
   Fermi sea is allowed with probability of the relative depletion of the Fermi sea



#### **High-Energy Fragments**

ISABEL was used <u>without</u> the additional coalescence model

No attempt was made to predict the production of high-enegy "heavy" fragments



## Typical Results ISABEL + GEMINI





## <sup>208</sup>Pb(p,X)n, 63 MeV





## <sup>208</sup>Pb(p,X)p, 63 MeV





## <sup>208</sup>Pb(p,X)n, 256 MeV





## <sup>208</sup>Pb(p,X)π<sup>-</sup>, 730 MeV





## <sup>208</sup>Pb(p,X)π<sup>+</sup>, 730 MeV





## <sup>208</sup>Pb(p,X)n, 800 MeV





## <sup>208</sup>Pb(p,X)p, 800 MeV



## <sup>208</sup>Pb(p,X), 1000 MeV







## <sup>208</sup>Pb(p,X), 1000 MeV





## <sup>208</sup>Pb(p,X), 1000 MeV











## <sup>208</sup>Pb(p,X)n, 1600 MeV











#### <sup>63</sup>Cu(p,X)π<sup>+</sup>, 730 MeV





#### <sup>56</sup>Fe(n,X)n, 65 MeV





# <sup>62</sup>Fe(p,X)p, 62 MeV





#### <sup>62</sup>Fe(p,X), 300 MeV





#### <sup>62</sup>Fe(p,X), 300 MeV





## <sup>62</sup>Fe(p,X), 300 MeV





## <sup>62</sup>Fe(p,X)n, 800 MeV





## <sup>62</sup>Fe(p,X), 1000 MeV





#### <sup>62</sup>Fe(p,X), 1000 MeV





#### <sup>62</sup>Fe(p,X), 1000 MeV





## <sup>62</sup>Fe(p,X)n, 1200 MeV





## <sup>62</sup>Fe(p,X)n, 1600 MeV





## <sup>62</sup>Fe(p,X)n, 3000 MeV





## <sup>238</sup>U(p,X), 1000 MeV





## <sup>238</sup>U(p,X), 1000 MeV





# **Thank You!**

## **Questions, Remarks?**

